

## Global illumination COM3404

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## Global illumination

Illumination  $I(\mathbf{x}, \mathbf{x}')$  at  $\mathbf{x}$  from surface at  $\mathbf{x}'$  is sum of

- light emitted by surface at  $\mathbf{x}'$ , denoted  $\epsilon(\mathbf{x}, \mathbf{x}')$
- sum of light scattered from  $\mathbf{x}'$  towards  $\mathbf{x}$  from other surfaces  $\mathbf{x}''$

Mathematically:

$$I(\mathbf{x}, \mathbf{x}') = g(\mathbf{x}, \mathbf{x}') \left[ \epsilon(\mathbf{x}, \mathbf{x}') + \int_S \rho(\mathbf{x}, \mathbf{x}', \mathbf{x}'') I(\mathbf{x}', \mathbf{x}'') d\mathbf{x}'' \right]$$

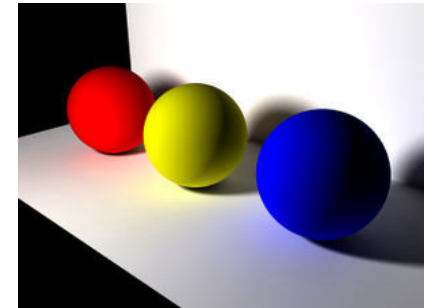
$g(\mathbf{x}, \mathbf{x}')$  *visibility function*. 0 if  $\mathbf{x}$  is not visible from  $\mathbf{x}'$ ; otherwise  $1/\text{dist}(\mathbf{x}, \mathbf{x}')^2$

$\rho(\mathbf{x}, \mathbf{x}', \mathbf{x}'')$  *bi-directional reflectance distribution function*, BRDF.

- Fraction of energy arriving at  $\mathbf{x}'$  from  $\mathbf{x}''$  that is scattered towards  $\mathbf{x}$ .
- Appearance depends on illumination (wavelength), the viewing direction, together with material properties
- Phong model is a simple example

## Outline

- 1 Global illumination
- 2 Whitted ray tracing
  - Recursion depth
  - Computation
- 3 Radiosity
  - Discrete approximation
  - Form factors



## References

- Fundamentals of 3D Computer Graphics. Watt. Chapters 10 & 12.
- Computer Graphics: Principles and Practice. Foley et al (1995).
- <http://freespace.virgin.net/hugo.elias/radiosity/radiosity.htm>
- "An empirical comparison of radiosity algorithms," A.J. Willmott and P.S. Heckbert, Carnegie Mellon Technical Report CMU-CS-97-115  
<http://www.cs.cmu.edu/~radiosity/emprad-tr.html>

## Rendering equation

### Components

- Model for light emitted by each surface,  $\epsilon()$
- A model for the BRDF  $\rho()$  of each surface
- A method of evaluating the visibility function

### Difficulties

**Complexity** Integral is so complex that it cannot be evaluated analytically.

- Possible numerical approximations by Monte-Carlo integration.
- Practical algorithms reduce complexity by approximation.

**View independence** Rendering equation is independent of the view.

- View dependent algorithms (eg. ray tracing, path tracing) reduce complexity by calculating illumination from a particular view.
- Radiosity method models view-independent diffuse illumination.

**Recursive** Expression for  $I(\mathbf{x}, \mathbf{x}')$  incorporates  $I(\mathbf{x}', \mathbf{x}'')$ . Leads to algorithms that recursively retrace the path of light arriving at a pixel in image plane: ray tracing, path tracing, distributed ray tracing.

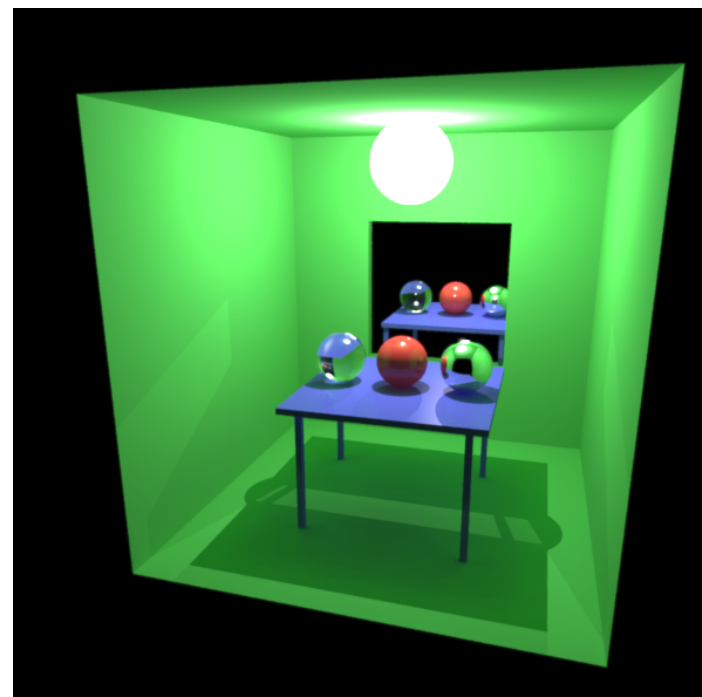
## Path notation

Four types of light transfer between diffuse and specularly reflecting surfaces:

- Diffuse – Diffuse** A diffuse surface reflects light onto a second diffuse surface
- Diffuse – Specular** A diffuse surface reflects light onto a specular surface
- Specular – Diffuse** A specular surface reflects light onto a diffuse surface
- Specular – Specular** A specular surface reflects light onto a second specular surface

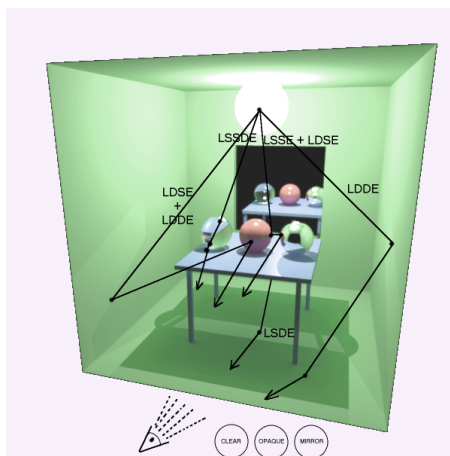
Categorise all paths from light **L** to eye **E** as **L(D|S)\*E**

Local reflection models permit LD|SE only



## Path notation

- LDDE** Shadow of table: light – wall – floor – eye
- LSE + LDE** Top of opaque ball: light – (specular|diffuse) – eye
- LDSE + LDDE** Bottom of opaque ball: light – wall – (specular|diffuse) – eye
- LSDE** Lighter shadow below table: light – mirror – floor – eye
- LSSDE** Caustic on table top: light – refraction through transparent sphere, S – refraction emerging from sphere, S – reflection at table top – eye



Clear, opaque and mirror balls on a table

## Whitted ray tracing

### Approximation

- Treat lights as *point* sources
- $\rho$  becomes law for perfect reflection and refraction
- Integral over scene reduced to sum over the path of a single ray



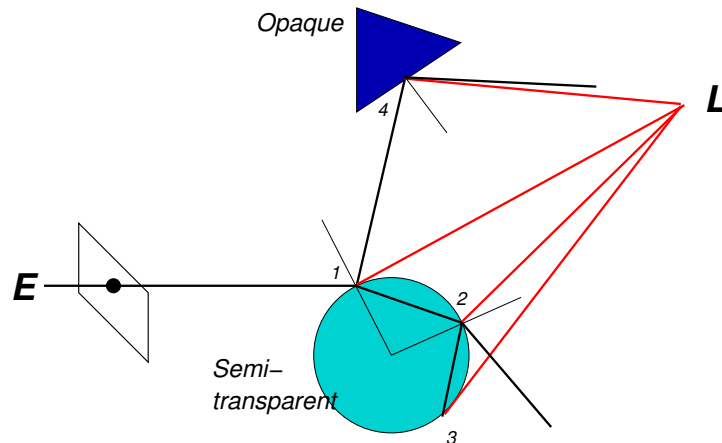
### View dependent

Calculation depends on the particular view

## Whitted ray tracing

### Trace a ray back from each pixel

At each surface: reflected and transmitted rays recursively traced back.  
Local specular and diffuse reflection from lights



## Whitted ray tracing

1

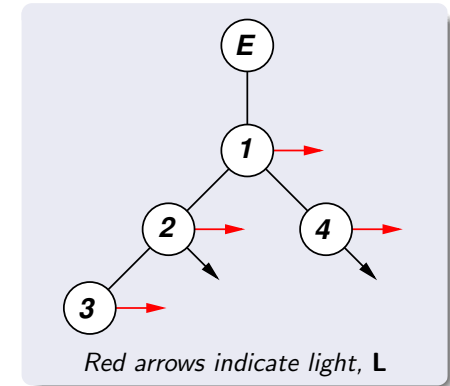
- Reflected ray from triangle traced back
- Refracted ray emerging from 2 traced back
- Local specular and diffuse reflection from L

2

- Reflected ray from 3 traced back
- Refracted ray traced, but hits nothing
- Local specular illumination from L

3

- Local specular illumination



4

- Reflected ray traced, but hits nothing
- No refraction/transmission
- Local specular/diffuse

## Whitted ray tracing

### Global model

*Specular* reflections from every surface computed and traced

### Local model

- Diffuse and specular reflection from light sources computed at every surface
- But diffuse reflections not propagated

### Opaque and diffuse spheres

LSSE Global: Light-Opaque-Mirror-Eye

LDSE Local: Light-Opaque-Mirror-Eye

LSE Global: Light-Mirror-Eye

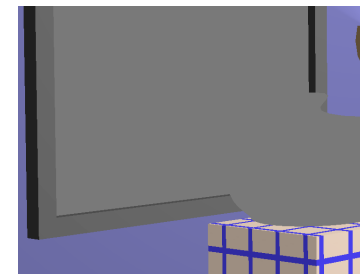
LDS\*E + LS\*E paths modelled



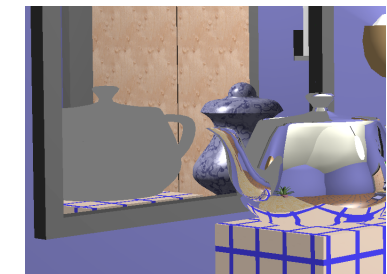
## Recursion depth

### Recursion

- Stopped when
- ray hits perfectly diffuse surface
  - rays meets scene boundary or bounding sphere.
  - recursion depth too great

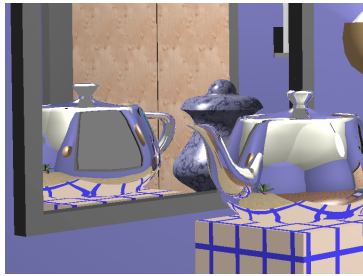


Depth 0

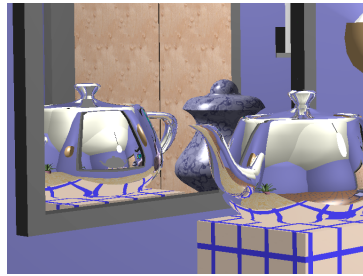


Depth 1

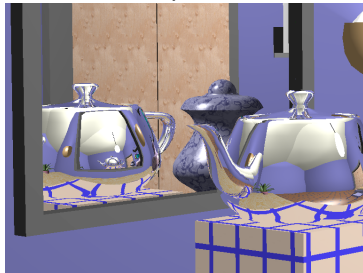
## Recursion depth



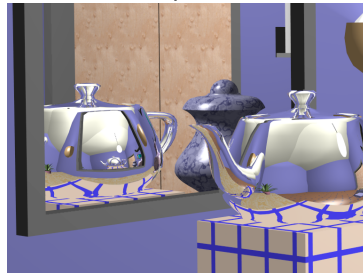
Depth 2



Depth 3



Depth 2

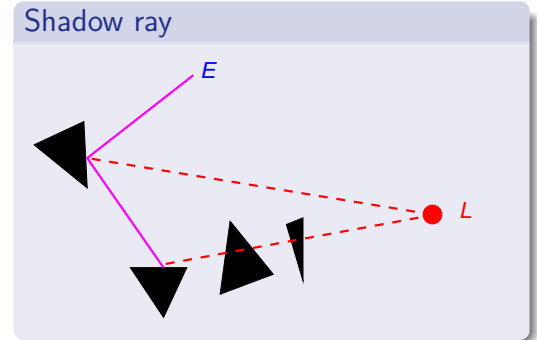


Depth 7 anti-aliased

## Shadows

Since light sources are point sources, only perfect, hard-edged shadows – the umbra – are calculated.

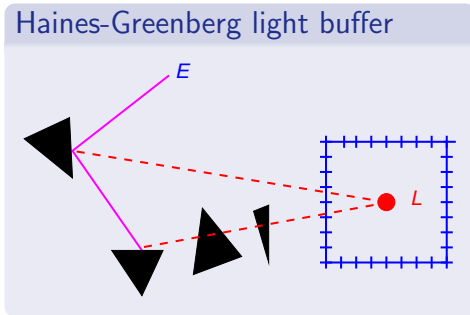
Calculate shadows by reducing illumination at a intersection  $\mathbf{x}$  if *shadow ray* from  $\mathbf{x}$  to light intersects another object.



## Shadows: Haines-Greenberg buffer

Intersection tests are expensive.

- Does not use coherence
- Reduced by storing polygons that intersect light rays in Haines-Greenberg light buffer.
- Each cell contains a list of polygons/objects and depths intersected by a ray leaving the light through the cell.
- When shadow testing, keep the opaque object which shadowed each light for each ray tree node. Check this object first at that node on the next shadow test as it is likely to be the shadowing object again.
- Do not calculate surface normals etc until it is verified the object is not in shadow.



## Computation

Recursive algorithm calculating the illumination at  $\mathbf{x}$  as

$$I(\mathbf{x}) = I_{local}(\mathbf{x}) + k_{rg}I(\mathbf{x}_r) + k_{tg}I(\mathbf{x}_t)$$

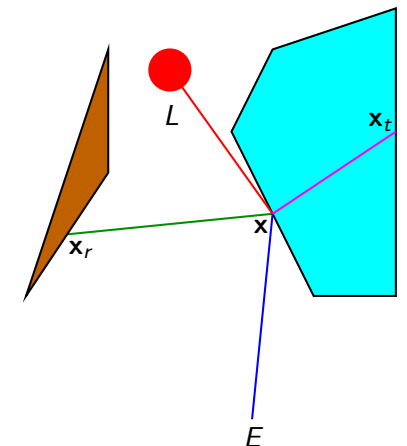
$I_{local}(\mathbf{x})$  Specular and diffuse reflection from lights

$\mathbf{x}_r$  Hit point discovered by tracing reflected ray from  $\mathbf{x}$

$\mathbf{x}_t$  Hit point discovered by tracing transmitted ray from  $\mathbf{x}$

$k_{rg}$  Global reflection coefficient

$k_{tg}$  Global transmission coefficient



### ShootRay(ray)

- intersection test
- if ray intersects an object
  - get normal at intersection
  - calculate local intensity  $I_{local}$
  - depth--
  - if depth > 0
    - calculate and shoot reflected ray
    - calculate and shoot refracted ray

### calculate and shoot reflected ray

- calculate reflection vector; store in reflected-ray structure
- RayOrigin := intersection point
- Attenuate ray: multiply  $k_{rg}$  by previous value
- ShootRay(reflected-ray)
- if reflected-ray intersects an object
  - combine colours  $k_{rg}I$  with  $I_{local}$

### calculate and shoot refracted ray

- if ray is entering object
  - accumulate refractive index
  - increment number of objects ray is currently inside
  - calculate refraction vector in refracted-ray
- else
  - de-accumulate refractive index
  - decrement number of objects ray is currently inside
  - calculate refraction vector in refracted-ray
- RayOrigin := intersection point
- Attenuate ray by  $k_{tg}$
- if refracted-ray intersects an object
  - combine colours  $k_{tg}I$  with  $I_{local}$

### Adaptive depth control

If product of transmission and reflection coefficients is small, stop tracing as contribution of ray is small.

### Eliminate first hit intersection calculations

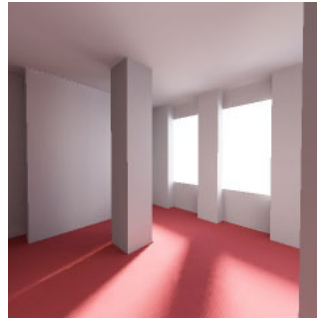
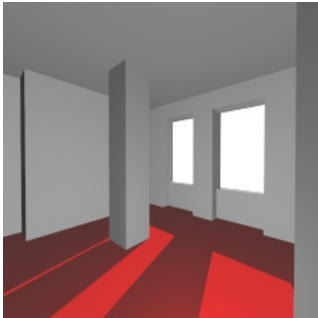
In many scenes rays make only one or two intersections, so using a modified Z-buffer algorithm pre-compute first intersection objects for each pixel in image plane.

### Bound objects with simple shapes

Only make detailed intersection calculations if ray is known to enter a simple bounding volume.

**Octrees** allow rapid location of objects close to the ray's path so intersections need only be calculated for these likely objects. Exploits spatial coherence of objects.

## Radiosity



### Ray traced

- Spot light outside window, ambient light
- Surfaces are diffuse reflectors
- Some faces invisible
- Hard shadows

### Radiosity

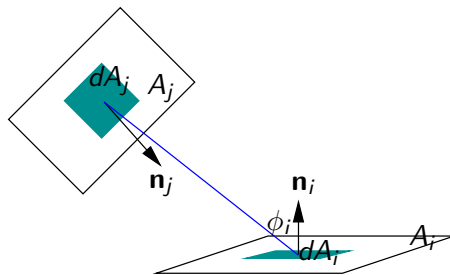
- Light source is image of the sky outside room
- Entire room lit and visible
- Soft shadows
- Graduated changes in light across faces
- Colours "bleed" across surfaces

## Form factor

$$F_{ij} = \frac{\text{Radiative energy leaving } A_i \text{ that arrives at } A_j}{\text{Radiative energy leaving } A_i \text{ in all directions}}$$

$$F_{ij} = \frac{1}{A_i} \int_{A_i} \int_{A_j} \frac{\cos \phi_i \cos \phi_j}{\pi r^2} dA_i dA_j$$

- Should be modified with the visibility function if patches are obscured
- Function of geometry; computed once for stationary scenes



## Radiosity

Radiosity: energy per unit area leaving a patch per unit time

- Sum of emitted and reflected energy
- For  $i$ th patch

$$B_i dA_i = E_i dA_i + R_i \int_j B_j F_{ji} dA_j$$

radiosity  $\times$  area = emitted + reflected from all other patches

$B_i$  radiosity of patch  $i$

$dA_i$  area of patch  $i$

$E_i$  emitted energy per unit area

$R_i$  reflectance of patch  $i$

$F_{ij}$  form factor: expresses geometric relationship between patches  $i$  and  $j$

## Discrete approximation

$$B_i dA_i = E_i dA_i + R_i \int_j B_j F_{ji} dA_j$$

- $F_{ij} dA_i = F_{ji} dA_j$
- Approximate integral by sum over discrete patches

$$B_i = E_i + R_i \sum_{j \neq i} B_j F_{ij}$$

In matrix form:

$$\begin{bmatrix} 1 - R_1 F_{11} & R_1 F_{12} & \dots & -R_1 F_{1N} \\ 1 - R_2 F_{21} & R_2 F_{22} & \dots & -R_2 F_{2N} \\ \vdots & \vdots & \ddots & \vdots \\ 1 - R_N F_{N1} & R_N F_{N2} & \dots & -R_N F_{NN} \end{bmatrix} \begin{bmatrix} B_1 \\ B_2 \\ \vdots \\ B_N \end{bmatrix} = \begin{bmatrix} E_1 \\ E_2 \\ \vdots \\ E_N \end{bmatrix}$$

## Radiosity method

$$\begin{bmatrix} 1 - R_1 F_{11} & R_1 F_{12} & \dots & -R_1 F_{1N} \\ 1 - R_2 F_{21} & R_2 F_{22} & \dots & -R_2 F_{2N} \\ \vdots & \vdots & \ddots & \vdots \\ 1 - R_N F_{N1} & R_N F_{N2} & \dots & -R_N F_{NN} \end{bmatrix} \begin{bmatrix} B_1 \\ B_2 \\ \vdots \\ B_N \end{bmatrix} = \begin{bmatrix} E_1 \\ E_2 \\ \vdots \\ E_N \end{bmatrix}$$

$\mathbf{RB} = \mathbf{E}$

- Emissions  $E_i$  are defined by the model  
Only non-zero where for light sources
- Elements of  $\mathbf{R}$  are defined by the model and BDRF for this wavelength
- Radiosities are solution to equations:  $\mathbf{B} = \mathbf{R}^{-1}\mathbf{E}$
- Separate system of equations for each wavelength or colour of light
- After radiosities are computed patches are rendered with Gouraud/Phong shading to interpolate across patches

## Gathering and shooting

Repeat until convergence:

for each patch  $i$ :

**Gather** Contribution of patch  $j$  to patch  $i$  is

$$B_i = E_i + R_i \sum_{j \neq i} B_j F_{ij}$$

so new estimate at iteration  $k + 1$  is

$$B_i^{(k+1)} = E_i + R_i \sum_{j=1}^N F_{ij} B_j^{(k)}$$

**Shoot** Contribution of patch  $i$  to  $B_j$  is  $R_j B_i F_{ji}$   
Update all patches  $j$  with change due to change in radiosity  
at  $i$ th patch,  $\Delta B_i = B_i^{(k+1)} - B_i^{(k)}$

$$B_j^{(k+1)} = B_j^{(k)} + R_j F_{ji} \Delta B_i$$

## Progressive refinement: gathering and shooting

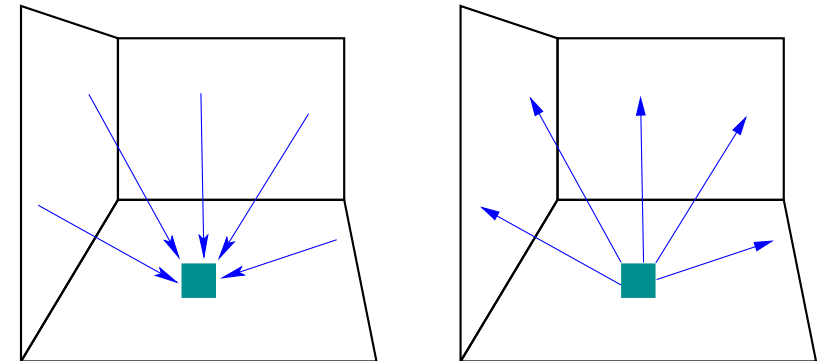
- Direct solution of  $\mathbf{B} = \mathbf{R}^{-1}\mathbf{E}$  is expensive in space and time
- Efficient solution by successive refinements of an approximate solution

Repeat until convergence:

for each patch:

**Gather** contributions for patch  $i$  from all other patches

**Shoot** radiosity from patch  $i$  to all other patches

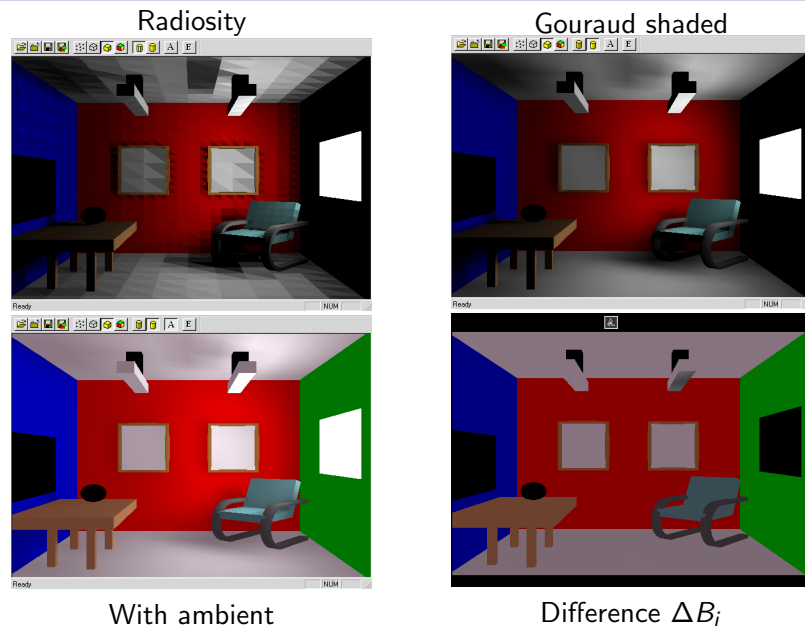


## Gathering and shooting

- Treat patches in order of emitting patches first, followed by those that have received a lot of light, etc.
- In the early stages of solution ambient light can be added to make darker parts of the solution visible. Ambient contribution removed as solution is refined.



## Example: 20 iterations

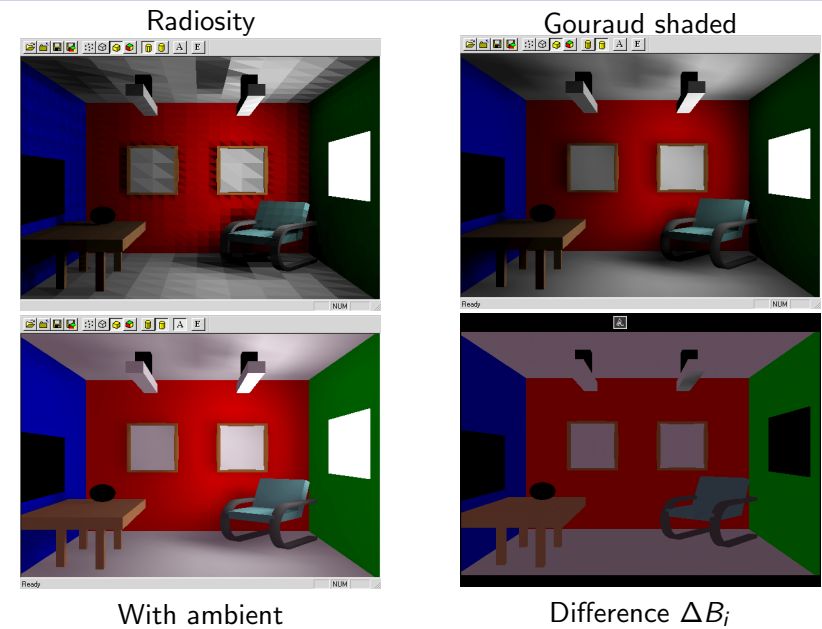


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## Example: 250 iterations

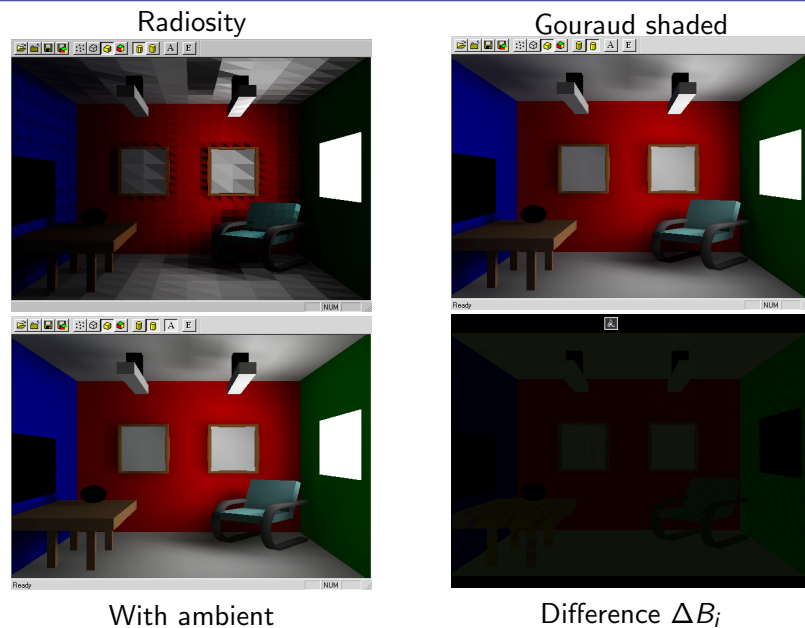


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## Example: 5000 iterations

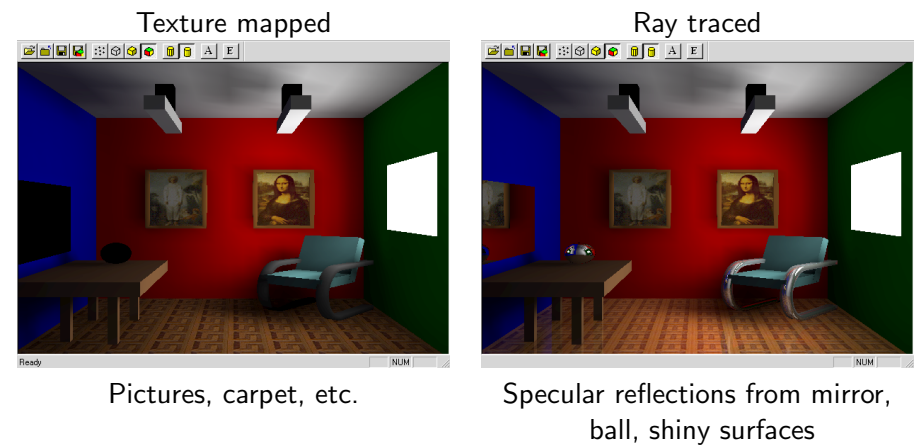


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## Example:



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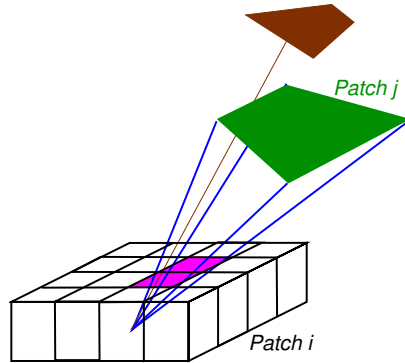
$$F_{ij} = \frac{\text{Radiative energy leaving } A_i \text{ that arrives at } A_j}{\text{Radiative energy leaving } A_i \text{ in all directions}} \approx \int_{A_j} \frac{\cos \phi_i \cos \phi_j}{\pi r^2} dA_j$$

if  $A_i$  is approximated by a small patch at its centre.

- Calculate  $F_{ij}$  by projecting patch  $j$  onto a *hemisphere* of pixels centred on patch  $i$
- Form factor for each hemisphere pixel

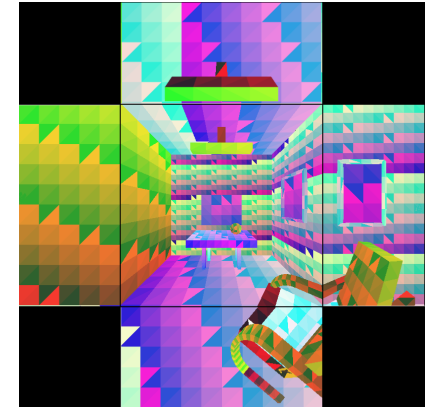
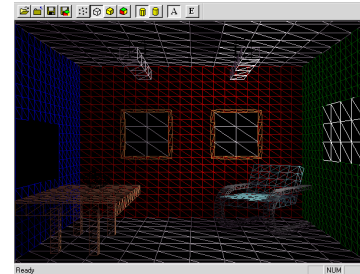
$$\Delta F_a = \frac{\cos \phi_i \cos \phi_j}{\pi r^2} \Delta A_a$$

- Projections calculated by modified Z-buffer algorithm



$$F_{ij} \approx \sum \Delta F_a$$

sum is over all pixels in the hemisphere to which patch  $j$  projects



Hemicube for a patch on the window

## Summary

### Global illumination method

- Handles diffuse light sources
- Renders shadows naturally
- View independent
- Form factors need only be calculated once for static scenes

### Disadvantages

- Computationally expensive, even with hemicube approximation
  - Many iterations required before scene emerges
  - Form factors must be recalculated for moving objects
- Artifacts may result from poor (too coarse) meshing
  - Fine meshes needed where light intensity changes rapidly
  - Adaptive schemes to refine mesh where gradient is high